## **Undergrad Exam Solution Fall 2019**

- a) Sol. Both linear and angular momenta are remain conserved before and after collision.
- **Sol.** The new center of mass after collision  $\frac{m_{\frac{L}{2}}^{L} + mL}{2m} = \frac{3L}{4}$ The moment of inertia of combined object (point mass + rod) around its center of mass can be evaluated as;  $I_{cm} = (m_{\frac{L^2}{12}}^{L^2} + m_{\frac{L^2}{16}}^{L^2}) + m_{\frac{L^2}{16}}^{L^2} = > = m_{\frac{L^2}{12}}^{L^2} + m_{\frac{L^2}{8}}^{L^2} = \frac{5mL^2}{24}$ .
- c) Sol. By law of conservation of linear momentum

$$P_i = P_f$$

 $mv_0=2mv_{cm}=>v_{cm}=v_0/2$  (linear center of mass velocity)

And angular momentum conservation law

$$L_i = L_f$$

$$mv_0\frac{L}{4} = I_{cm}\omega$$

$$\mathrm{m}v_0\frac{L}{4} = \frac{5mL^2}{24}\omega$$
 where  $I_{cm} = \frac{5mL^2}{24}$ 

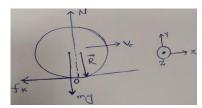
 $\omega = \frac{6v_0}{5L}$  required angular velocity and its direction is out of page!

d) Sol. The kinetic energy of the system

K.E = 
$$\frac{1}{2}$$
(2m  $v_{cm}^2$ ) +  $\frac{1}{2}$ ( $\frac{5mL^2}{24}$ ) ( $\frac{6v_0}{5L}$ )<sup>2</sup>  
=>  $\frac{1}{2}$  m[2 ( $\frac{v_0}{2}$ )<sup>2</sup> +  $\frac{3}{10}$   $v_0^2$ ]  
=>  $\frac{1}{2}$  m [ $\frac{1}{2}$   $v_0^2$  +  $\frac{3}{10}$   $v_0^2$ ]  
K.E =  $\frac{2}{5}$   $mv_0^2$ 

\*

Q. 2 a) Sol. From figure, we can write



$$\vec{V}(t=0) = v_0 \hat{x}$$

$$\vec{\omega}(t=0)=-\omega_0\hat{z}$$

$$v_0 > \omega_0 R = > \vec{f}_k = -f_k \hat{\mathbf{x}}$$
 (at touching point to ground) where  $f_k = \mu_k . N$ 

b) Sol. 
$$\hat{y}$$
: N-mg=  $ma_y$  since  $a_y$ =0  
=> N-mg=0 => N=mg  
 $\hat{x}$ :  $-f_k$ = $ma_x$  =>  $-\mu$ mg= $ma_x$   
 $a_x$  =  $-\mu$ g  
 $\hat{z}$ :  $-Rf_k$ = $I\alpha$ =>  $-\mu$ Rmg =  $\frac{2}{5}mR^2\alpha$   $\alpha$ = $-\frac{5\mu}{2R}g$ 

Center of mass speed;

$$\hat{v}_x(t) = (v_0 - \mu g t) \hat{x}$$

$$\vec{V}(t) = \vec{\omega}(t) \times \vec{R}$$

$$= > = (-\omega_0 - \frac{5\mu}{2R} g t) \hat{z} \times (-R) \hat{y}$$

$$= > = -(v_0 + \frac{5\mu}{2} g t) \hat{x}$$

Total speed of the body at touching point;

$$\hat{v}_{x}(t) + \vec{V}(t) = v_0 - \omega_0 R - \frac{7\mu}{2} gt$$
 (a)

At t=T, we have rolling without slipping, therefore, the total velocity is 0. Eq. (a) becomes

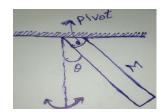
$$0 = v_0 - \omega_0 R - \frac{7\mu}{2} gT$$

$$= > v_0 - \omega_0 R - \frac{7\mu}{2} gT$$

$$T = \frac{2}{7} \frac{(v_0 - \omega_0 R)}{\mu g}$$

Q. 3 a) Sol. The moment of inertia of the rod at pivot point.

$$I_{cm} = \frac{1}{12}Ml^2 + \frac{1}{4}Ml^2$$



$$=> I_{cm} = \frac{1}{3} M l^2$$

The torque produced by the rod during its motion

$$\tau = r \times \vec{F} = \tau = r \vec{F} \sin \theta$$

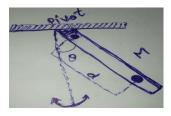
$$\tau = -\frac{l}{2} \text{Mgsin } \theta = \frac{1}{3} M l^2 \cdot \frac{d\theta^2}{dt^2}$$

For small deflection i.e.  $\sin \theta \approx \theta$ 

$$\frac{d\theta^2}{dt^2} + \frac{3g}{2l}\theta = 0. \implies \ddot{\theta} = \omega^2 \theta \qquad \omega = \sqrt{\frac{3g}{2l}}$$

Required equation of motion, which is similar to simple harmonic motion.

**b)** Sol. The total inertia of the rod is



$$I_{total} = \frac{1}{3}Ml^2 + md^2$$

Torque of the rod

$$\mathbf{\tau} = -\frac{l}{2} Mg \sin \theta - mg d \sin \theta$$

$$I_{total} \cdot \frac{d\theta^{2}}{dt^{2}} = -\frac{l}{2} Mg \sin \theta - mg dSin\theta \qquad \mathbf{\tau} = I_{total} \cdot \frac{d\theta^{2}}{dt^{2}}$$

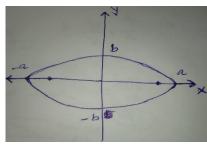
$$\tau = I_{total}. \frac{d\theta^2}{dt^2}$$

$$\Rightarrow \frac{d\theta^2}{dt^2} + \frac{Mgl/2 + mgd}{\frac{1}{3}Ml^2 + md^2}\theta = 0$$

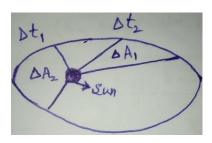
$$\Rightarrow \ddot{\theta} = \omega^2 \theta \qquad \omega = \sqrt{\frac{Mgl/2 + mgd}{\frac{1}{3}Ml^2 + md^2}}$$

$$\Rightarrow \frac{3g}{2l} = 2l = 3d = d = \frac{2l}{3}$$
 is the point at which the frequency of the oscillation remains the same.

**Q.4 a) Sol.** Equation 
$$\frac{x^2}{a^2} + \frac{y^2}{b^2} = 1$$



**b) Sol.**  $\frac{dA}{dt}$  = Constant due to the conservation of angular moment (Kepler's 2<sup>nd</sup> law)



i.e. 
$$\Delta A_1 = \Delta A_2$$
 if  $\Delta t_1 = \Delta t_2$ 

c) Sol. 
$$\frac{a^3}{T^2}$$
 = Constant (Kepler's 3<sup>rd</sup> law)

Proof. The centripetal force of the planet balance the gravitation

$$\vec{F}_c = \vec{F}_g$$

$$\frac{mv^2}{r} = \frac{GmM}{r^2}$$

$$v = \sqrt{\frac{GM}{r}}$$
Since  $v = r\omega = \frac{2\pi r}{T}$   $\omega = \frac{2\pi}{T}$ 

$$\frac{2\pi r}{T} = \sqrt{\frac{GM}{r}} = > (2\pi r/T)^2 = \frac{GM}{r}$$

$$T^2 = \frac{4\pi^2}{GM} r^3 = > \frac{r^3}{T^2} = \frac{GM}{4\pi^2} = constant$$